

# Diamond subwavelength gratings for mid-infrared AGPM coronagraph: manufacturing assessment

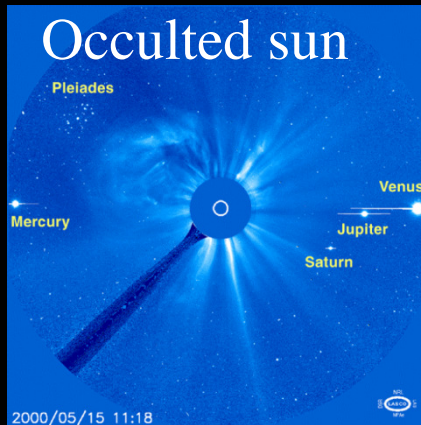
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PhD student

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P. Forsberg, M. Kuittinen and I. Vartianen

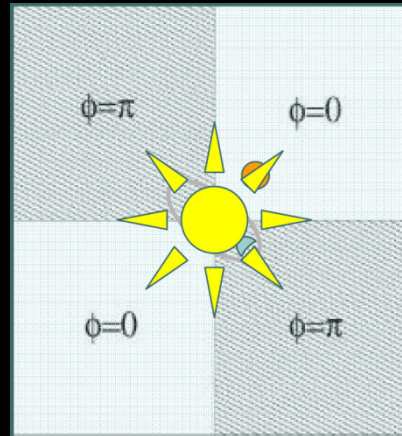
*EOS Annual Meeting, 26/10/2010*

# What is a phase mask coronagraph?

- Aim : direct imaging of high contrast astronomical scenes (e.g. exoplanets)
- Major constrains :
  - huge contrast ( $10^6$  in the thermal IR)
  - proximity :  $1''$  at 10 parsecs (Jupiter-like planets)
- Coronagraph = “hide” the star to unveil the hidden planet

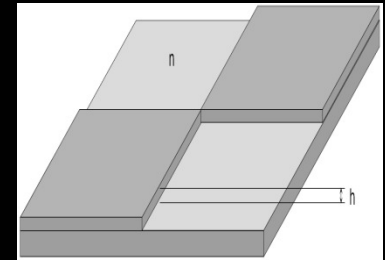


→ Amplitude mask



→ FQPM (*Four Quadrant Phase Mask*)

Star  
→ phase mask

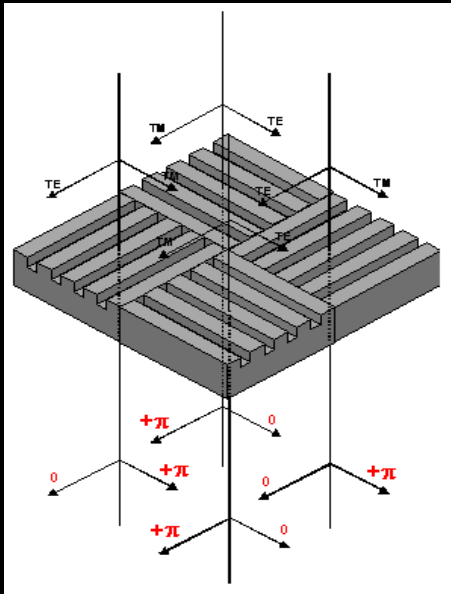


Transitions : *loss of information*

Chromaticity : *poor Signal to Noise ratio*

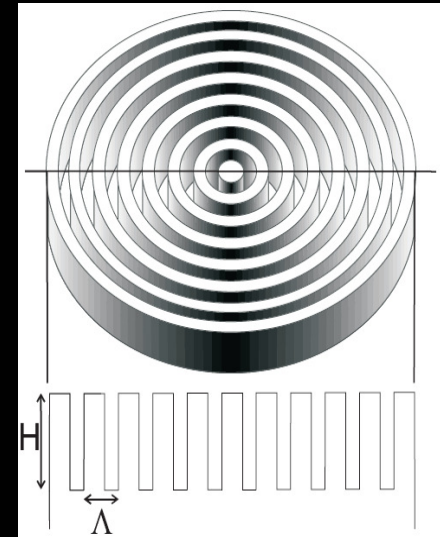
# The Annular Groove Phase Mask (AGPM)

- Proposed by Mawet et al. 2005
- Zeroth Order (sublambda => period <  $\lambda/n$ ) Grating (ZOG)
- Form birefringence  $\rightarrow \Delta\phi_{TE-TM} = \pi \rightarrow$  phase retarders
- Achromatic on wide spectral bands (in the visible or IR)

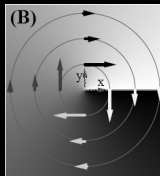


FQPM (*Four Quadrant Phase Mask*)

Vector Vortex Coronagraph (VVC)  
360° discovery space



AGPM (*Annular Groove Phase Mask*)

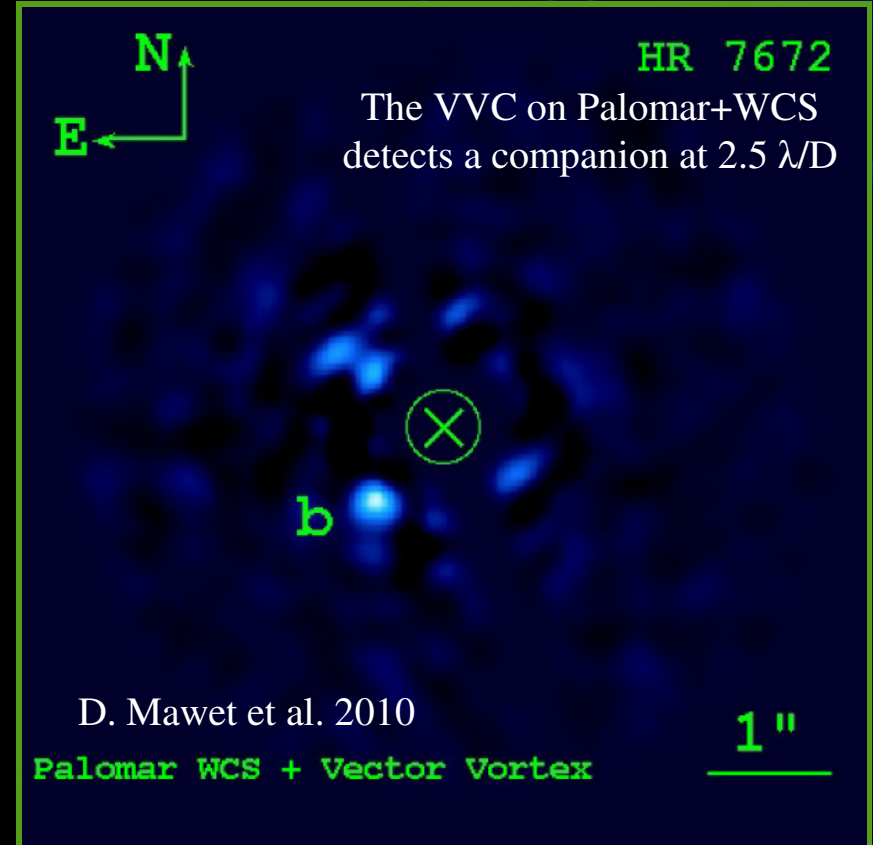


# Which spectral band ?

Other techniques for VVCs, easier to implement, exist: e.g. liquid-crystal polymers (LCP), Mawet et al. 2010

- ☺ lab demos in the visible and near-IR
- ☺ used on Palomar in the H- and K-bands
- ☹ technically limited to the K-band

In fact, these are limited to the visible and near-IR whereas the AGPM is suitable for any spectral band, from the visible to the thermal IR



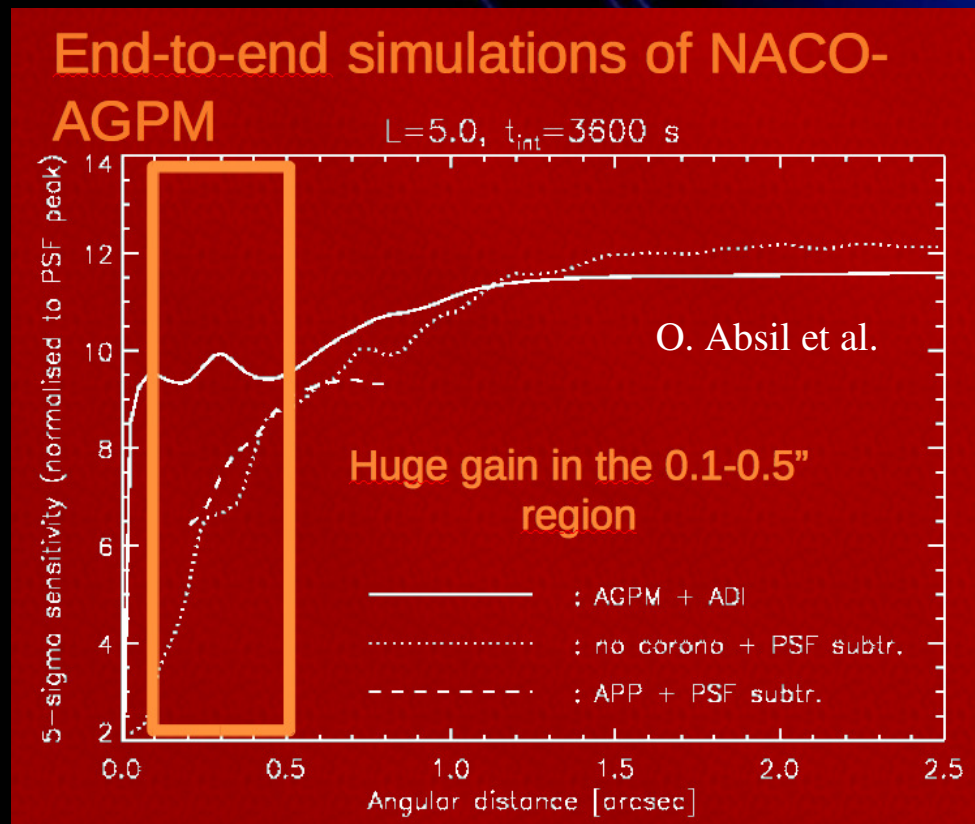
# Which spectral band ?

## L-band (3.5 – 4.1 $\mu\text{m}$ )

- recent success with NAOS-CONICA (Lagrange et al. 2010)
- important gain in the 0.1-0.5" region, compared to APP+PSF subtraction (see O. Absil's poster)

## N-band (8 – 13 $\mu\text{m}$ )

- the AGPM is foreseen for the upgrade of VISIR, and candidate for METIS on the future E-ELT
- subwavelength gratings are one of the only solutions at this wavelength

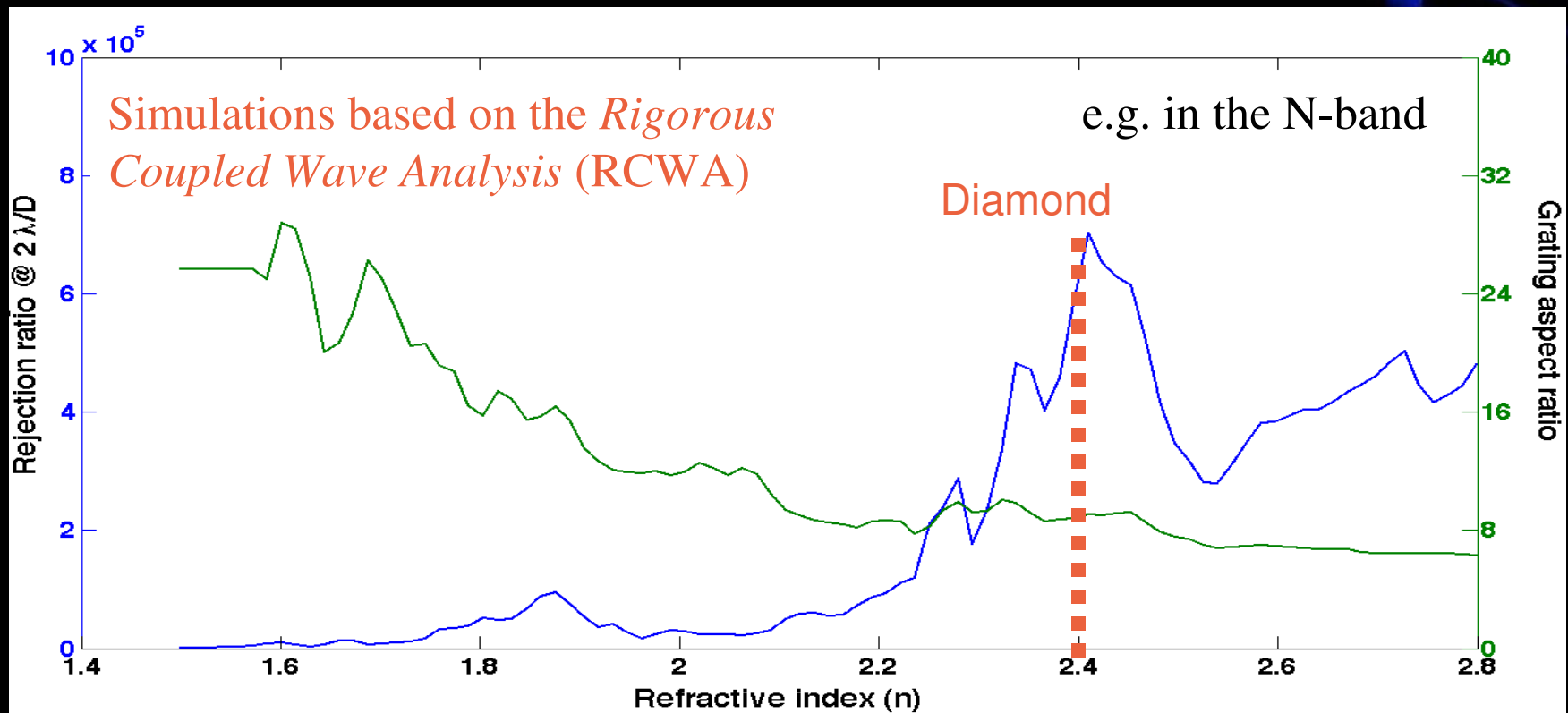
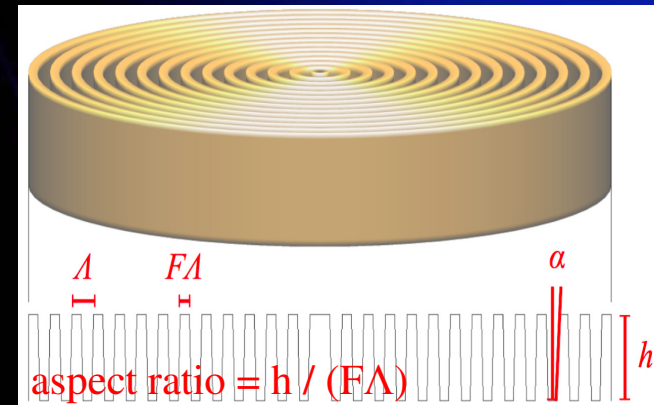




# Which substrate material ?

CVD diamond

- ☺ large spectrum : from visible to thermal-IR ( $\sim 20 \mu\text{m}$ )
- ☺ favorable mechanical and thermal properties
- ☺ high refractive index = shallow etching

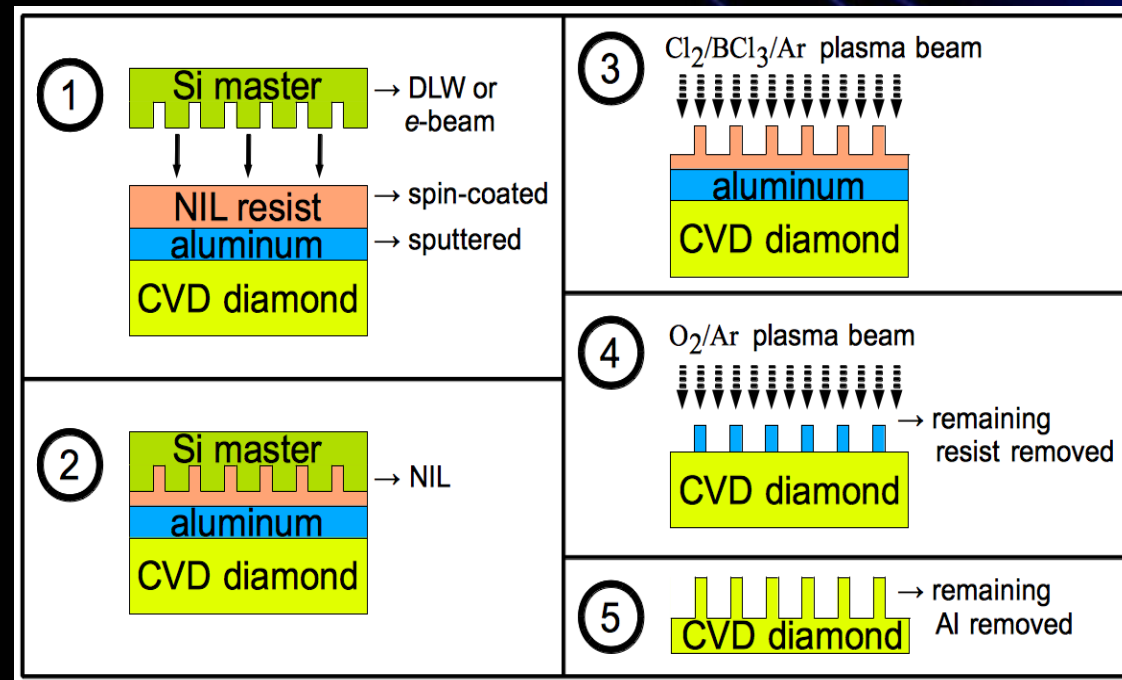


# First manufactured diamond AGPM

## Manufacturing process

Inter-University collaboration:

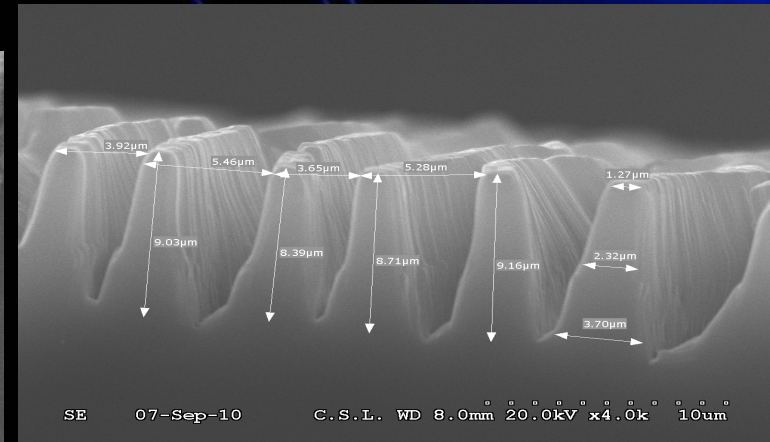
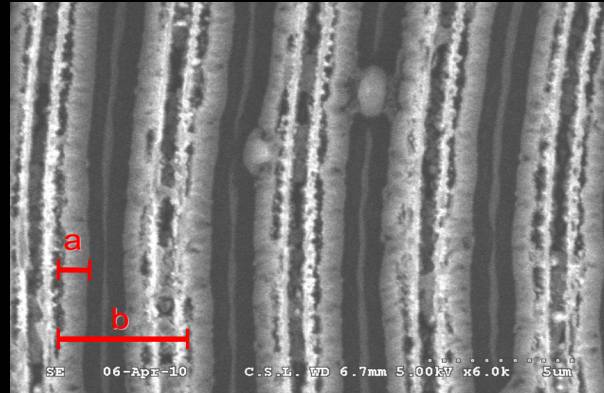
- e-beam mask at the University of Joensuu, Finland  
(M. Kuittinen and I. Vartanen)
- Nano-Imprint Lithography (NIL) and Reactive Ion Etching at the University of Uppsala, Sweden  
(M. Karlsson, F. Nikolajeff, P. Forsberg)
- Metrology and optical testing at the University of Liège, Belgium



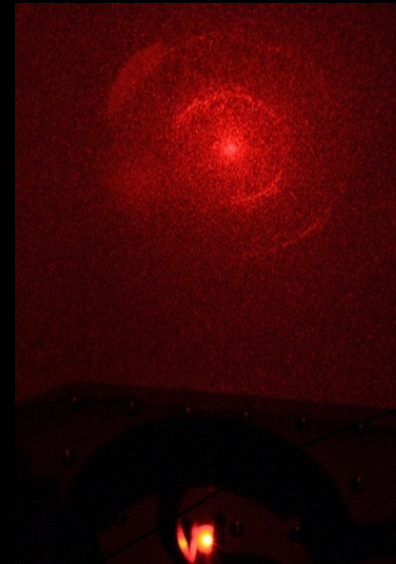
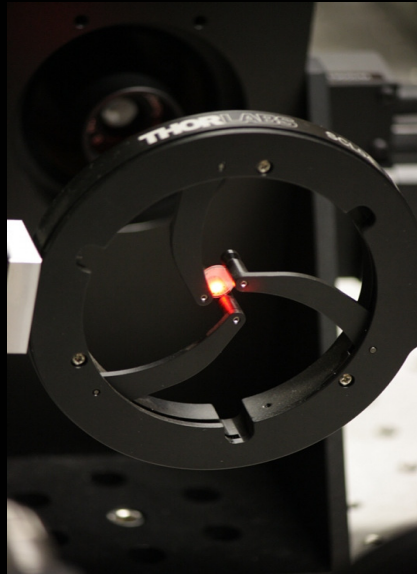
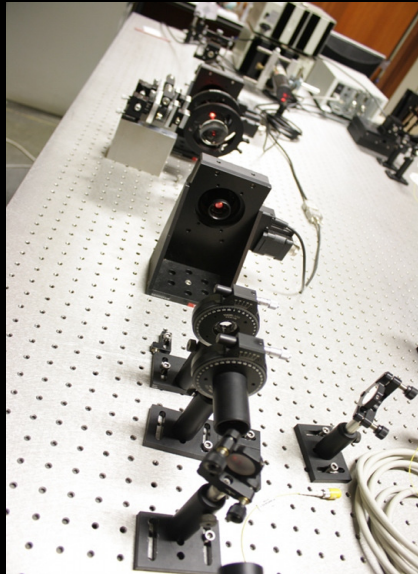
# First manufactured diamond AGPM

## Classical metrology + moulding

- interferometry
- SEM, AFM
- silicone moulding
- profile metrology



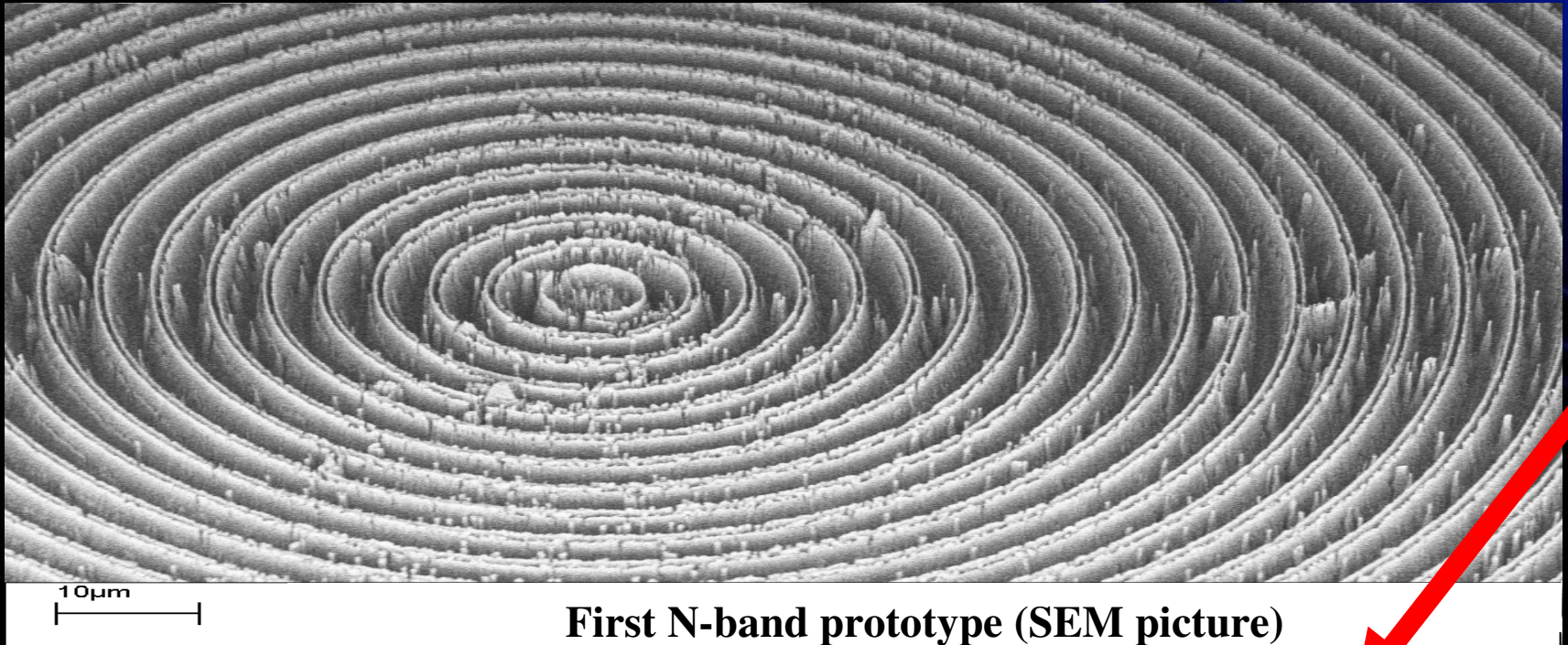
## Defractometry + scattering measurements



- optical bench
- HeNe laser (632.82 nm)
- 5 orders
- total integrated scattering (TIS) =  
< 0.4% (N-band)  
< 2.3% (L-band)



# First manufactured diamond AGPM



First N-band prototype (SEM picture)

- optimised for  $[10.5\mu\text{m} - 12.25\mu\text{m}]$
- performances simulated with RCWA  
(= mean Null Depth @  $2\lambda/D$ )
  - near the center:  $\mu \approx 10^{-5}$
  - away from the center:  $\mu \approx 10^{-3}$
- foreseen for the upgrade of VISIR

poor optical quality in the area away from the center, because of too small substrate (1cm diameter)

→ NIL process being improved  
→ bigger substrates (2cm diameter)

# Components currently being manufactured

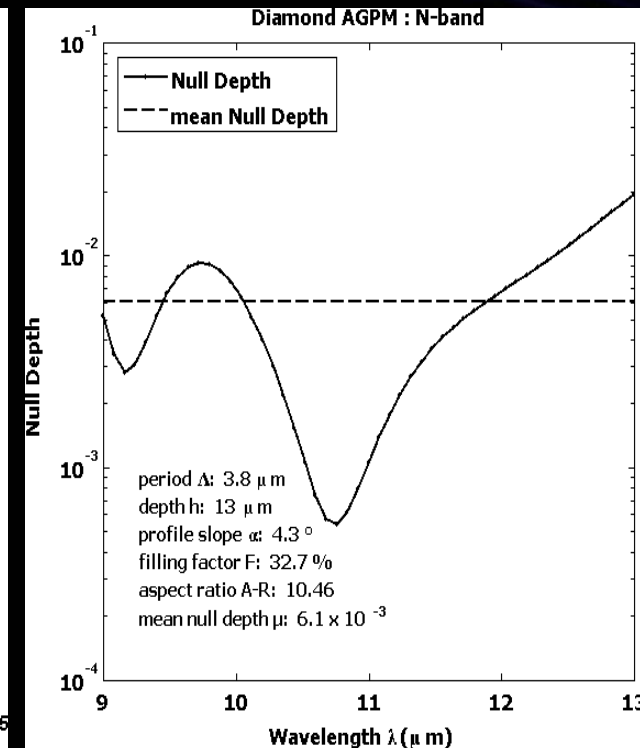
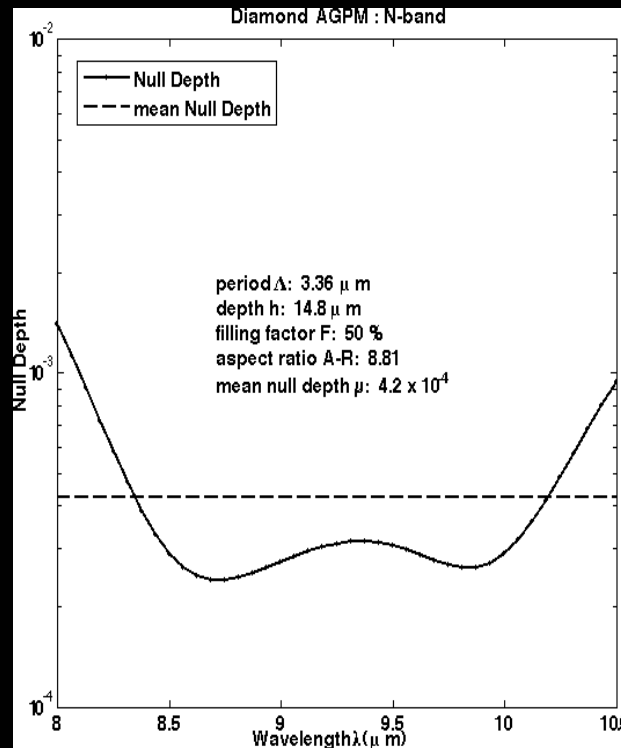
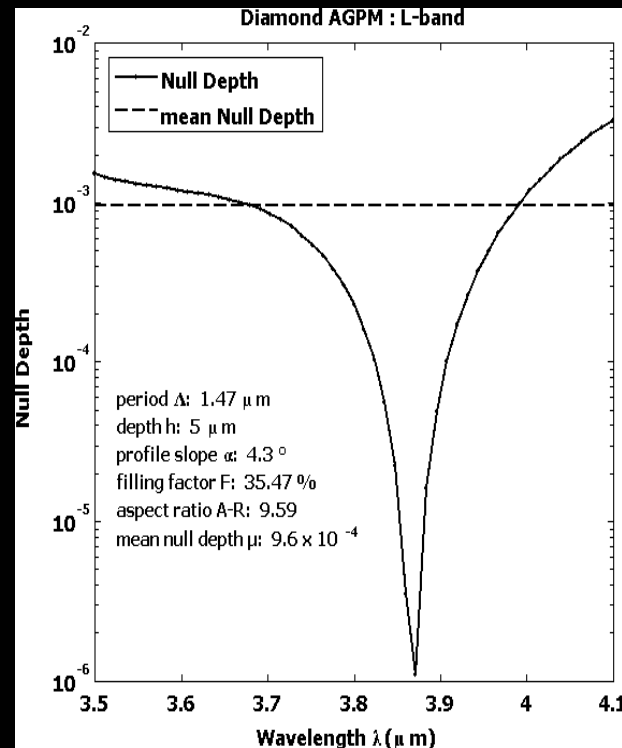
## L-band

- $[3.5 \mu\text{m} - 4.1 \mu\text{m}]$ :  $\mu \approx 5 \times 10^{-6}$  @  $2\lambda/D$

→ RCWA simulations :

## N-band : 2 components

- $[8 \mu\text{m} - 10.5 \mu\text{m}]$ :  $\mu \approx 2 \times 10^{-6}$  @  $2\lambda/D$
- $[9 \mu\text{m} - 13 \mu\text{m}]$ :  $\mu \approx 3 \times 10^{-5}$  @  $2\lambda/D$



# Summary

- The first diamond prototype has been manufactured
- Good metrology  $\rightarrow$  parameters well known
- Possible use on the sky with the future upgrade of VISIR at  $[10.5\mu\text{m} - 12.25\mu\text{m}]$  with an expected  $\mu \approx 10^{-3}-10^{-5}$  @  $2\lambda/D$
- Lessons learned, microfabrication techniques improved
- Next components are currently being manufactured with expected performances  $\mu < 10^{-5}$  @  $2\lambda/D$
- We focus on the L-band for NACO where the AGPM is very promising



Thank you for your attention !